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Geomagnetically Induced Currents (GICs) during space weather events in May and October 2024

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Geomagnetically induced currents (GICs)

GICs are induced by electric fields generated by any rapid changes in the magnetic field during various space weather events

$$\nabla \times \mathbf{E} = - \partial \mathbf{B} / \partial t$$

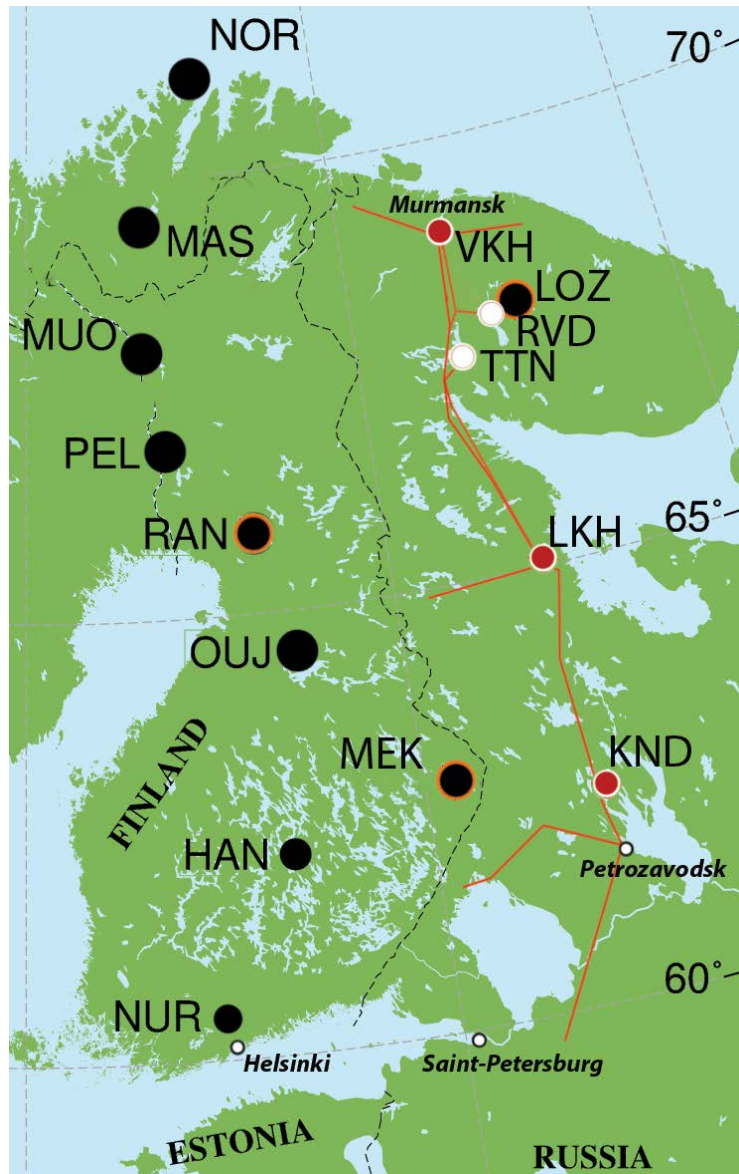
Such events include the development of a **sudden storm commencement** (SSC) pulse (Zhang *et al.*, 2015), a sharp increase in the **ring current** (Gil *et al.*, 2023), an intensification of **auroral electrojets** (Boteler, 1999; Despirak *et al.*, 2022), the generation of **low-frequency pulsations** (Apatenkov *et al.*, 2020; Yagova *et al.*, 2021; Setsko *et al.*, 2022).

The magnitude of GICs depends not only on the intensity of magnetic disturbances, but also on **ground conductivity** and the **configuration of the technological systems** (e.g. Viljanen and Pirjola, 2017; Love *et al.*, 2018a 2018b; Ingham *et al.*, 2023).

- The conductivity values for the region where the Karelian-Kola power line is located are almost the same.
- a direct comparison of GIC is not possible because transmission lines undergo outages, upgrades, or the addition or replacement of transformers.



- **direct GIC`s measurements in Russian electrical circuits**



Geographic coordinates:

Vykhodnoy (VKH) (68.8° N, 33.1° E),
Loukhi (LKH) (66.08° N, 33.12° E),
Kondopoga (KND) (62.2° N, 34.3° E)

System for monitor GICs in the existing Karelian-Kola power transmission line with a length of over 800 km (<http://eurisgic.ru/>).

“Vykhodnoy”, “Revda”, “Titan” (Murmansk region),
“Loukhi” and “Kondopoga” (Republic of Karelia)

The power transmission line stretched from south to north. According to Faraday's law, the variations of the Y-component of the geomagnetic field play a major role in GIC generation (Setsko et al., 2023; Selivanov et al., 2023).

VKH is a node station that collects GIC, according to Kirchhoff's law, from both the East-West and North-South lines, so the GIC at VKH is the sum of the currents induced by variations in the X- and Y-components of the magnetic field.

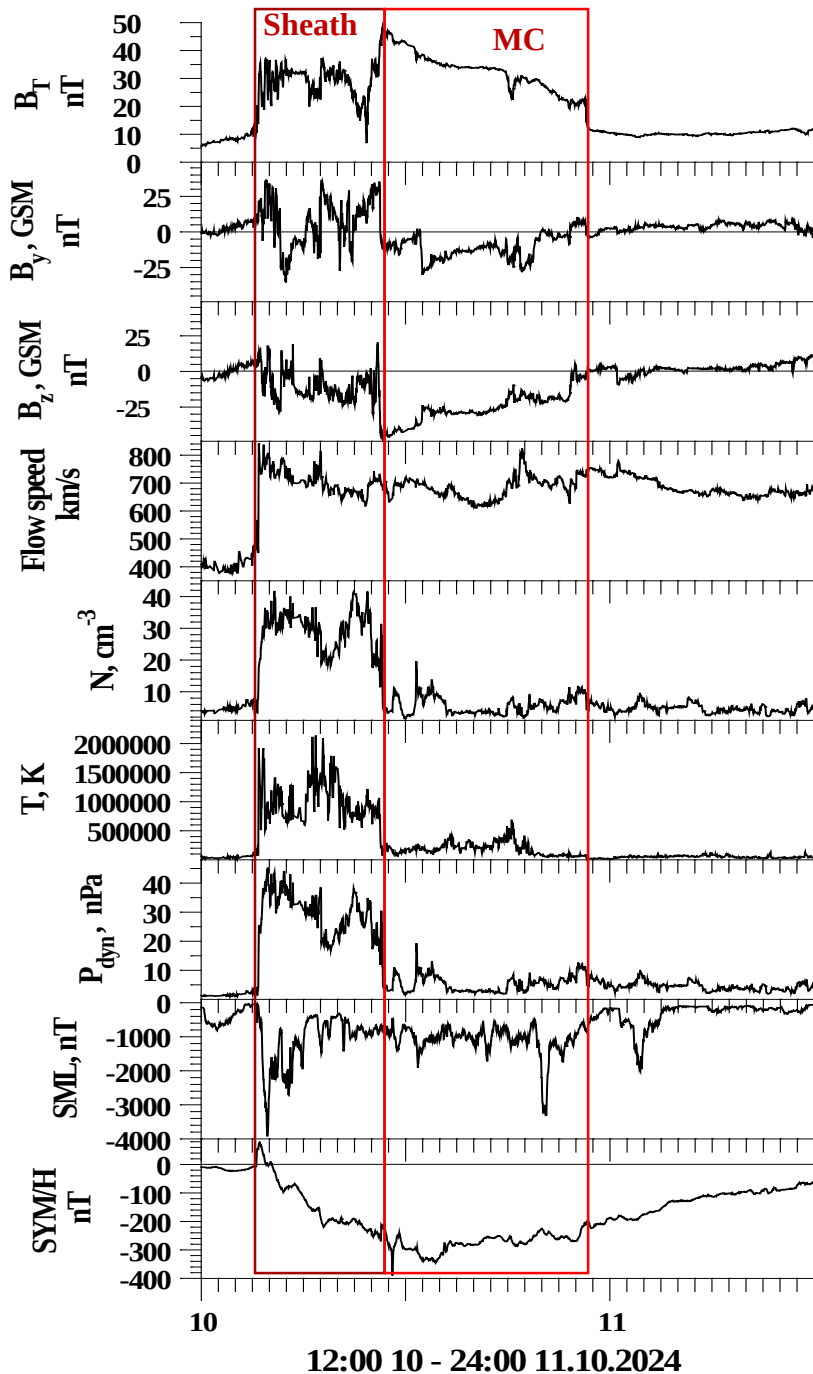
VKH station has two autotransformers, and an additional transformer nearby, i.e. three autotransformers for VKH node.

LKH station has two autotransformers,

KND station has only one.

Since the GIC is measured on a single grounded line of one autotransformer at each station, the total current in each node can be estimated by multiplying the measured value by the number of grounding points (1 at KND, 2 at LKH and 3 at VKH).

Solar wind and IMF conditions



Large magnetic storm ($\text{Dst} = -333 \text{ nT}$).

The differences from the May 2024 storm are;
the color of aurora (predominantly red)
the phase of magnetic storm (early main phase just after the sudden commencement).

The origins of the storm are the disk-center X1.8 flare at $\sim 02:00 \text{ UT}$ on October 9, 2024 and the full-halo coronal mass ejection (CME) with an initial speed of $\sim 1000 \text{ km/s}$. (*Kataoka et al., 2025*)

The interplanetary shock was observed at $\sim 15:00 \text{ UT}$ on October 10, 2024.

Just after the shock arrival or sudden commencement at the Earth at $\sim 15:20 \text{ UT}$,

an extremely large substorm occurred, so-called supersubstorm (SSS), $\sim 4000 \text{ nT}$ at $\sim 16 \text{ UT}$.

This is associated with a strong southward directing B_z of -20 nT , a high solar wind density of $30/\text{cc}$, and a high solar wind speed of $\sim 700 \text{ km/s}$

Several intense substorms $\sim 17:00$, $\sim 18:30$, $\sim 20:30$, $\sim 22:30 \text{ UT}$ were recorded after SSS, during Sheath

In the morning sector, during the magnetic cloud (MC), there were also several substorms $\sim 00:40$, $02:30$, $03:30$, $05:50 \text{ UT}$

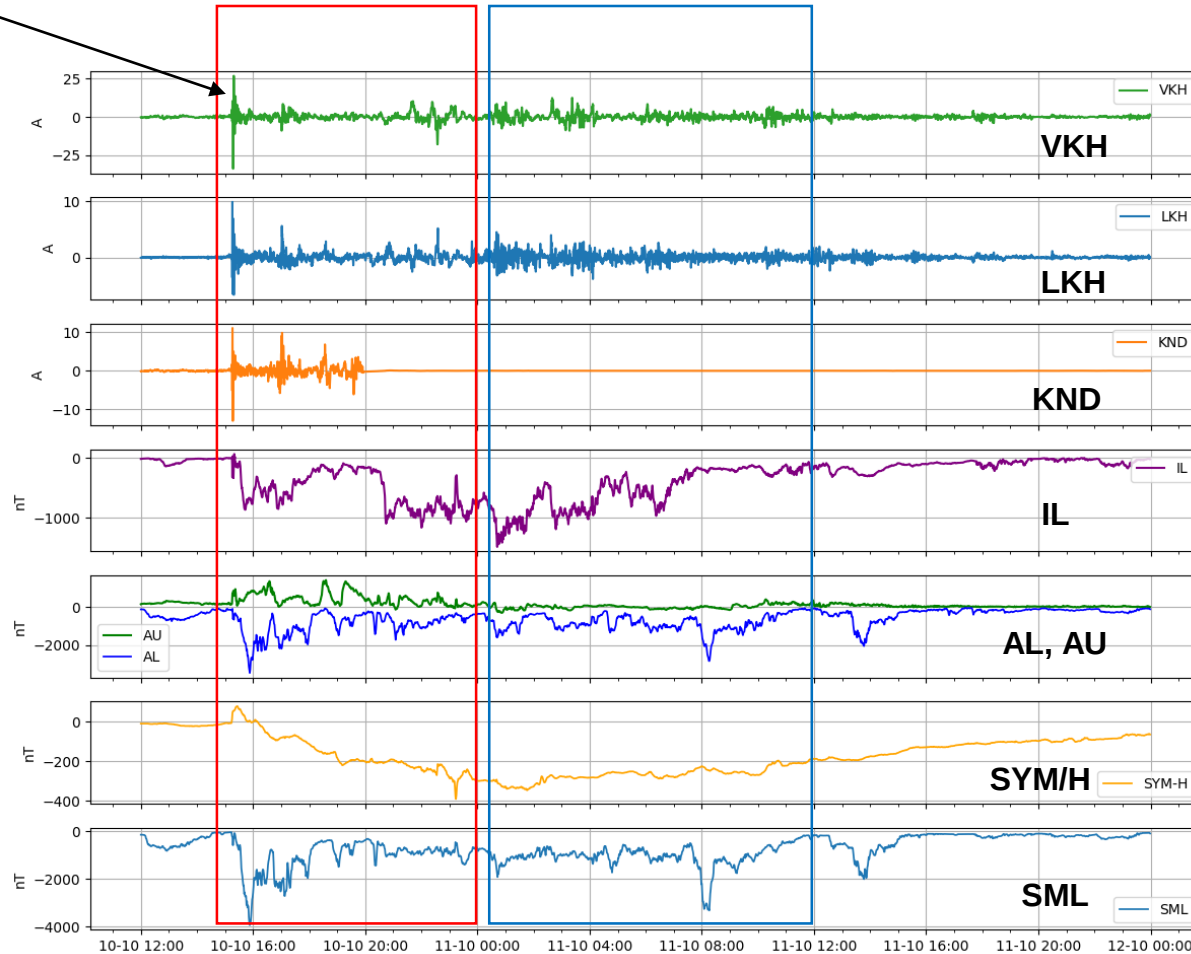
GICs and geomagnetic indexes

~15:20 UT,
~30A (VKH)
~10A (LKH)
~12A (KND)

~17:00 UT,
~7A (VKH)
~6A (LKH)
~10A (KND)

~22:40 UT,
~12 A (VKH)
~5A (LKH)

First interval Second interval



~00:40 UT,
~10A (VKH)
~4A (LKH)

~02:40 UT,
~12A (VKH)
~4A (LKH)

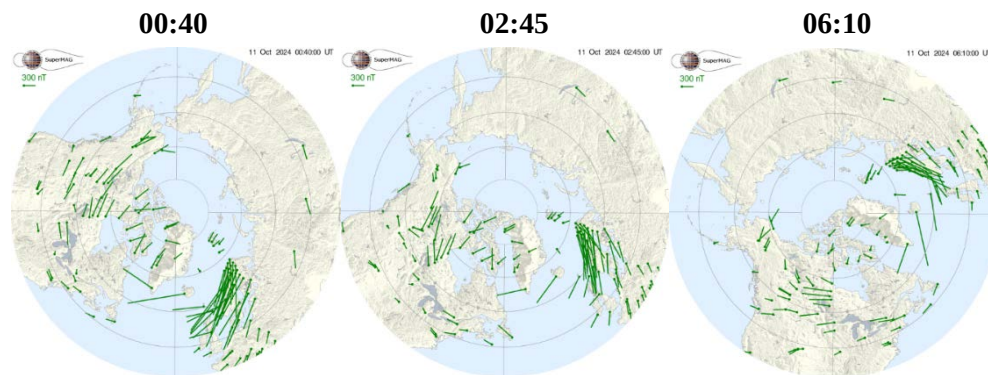
~06:00 UT,
~8A (VKH)
~3A (LKH)

- GIC sources in the Second interval:

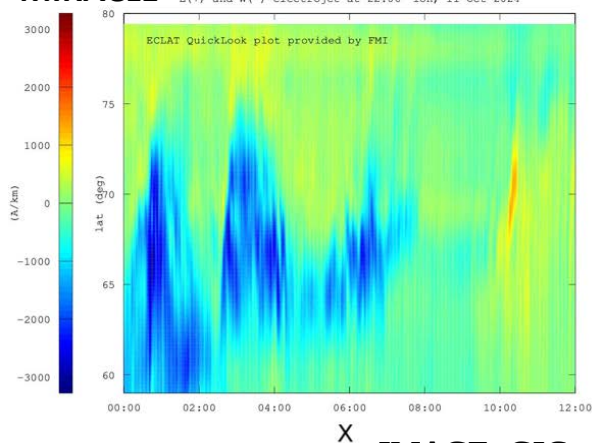
Pc5/Pi3 geomagnetic pulsations and substorms at the morning sector

Second interval

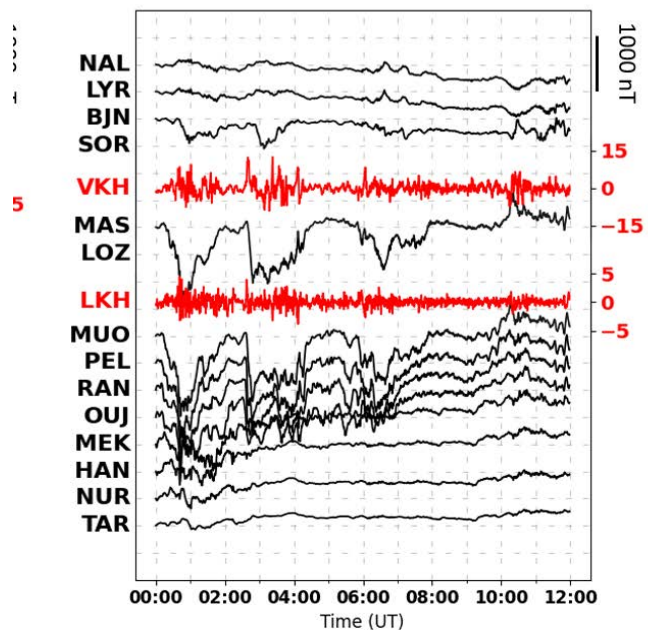
SuperMAG



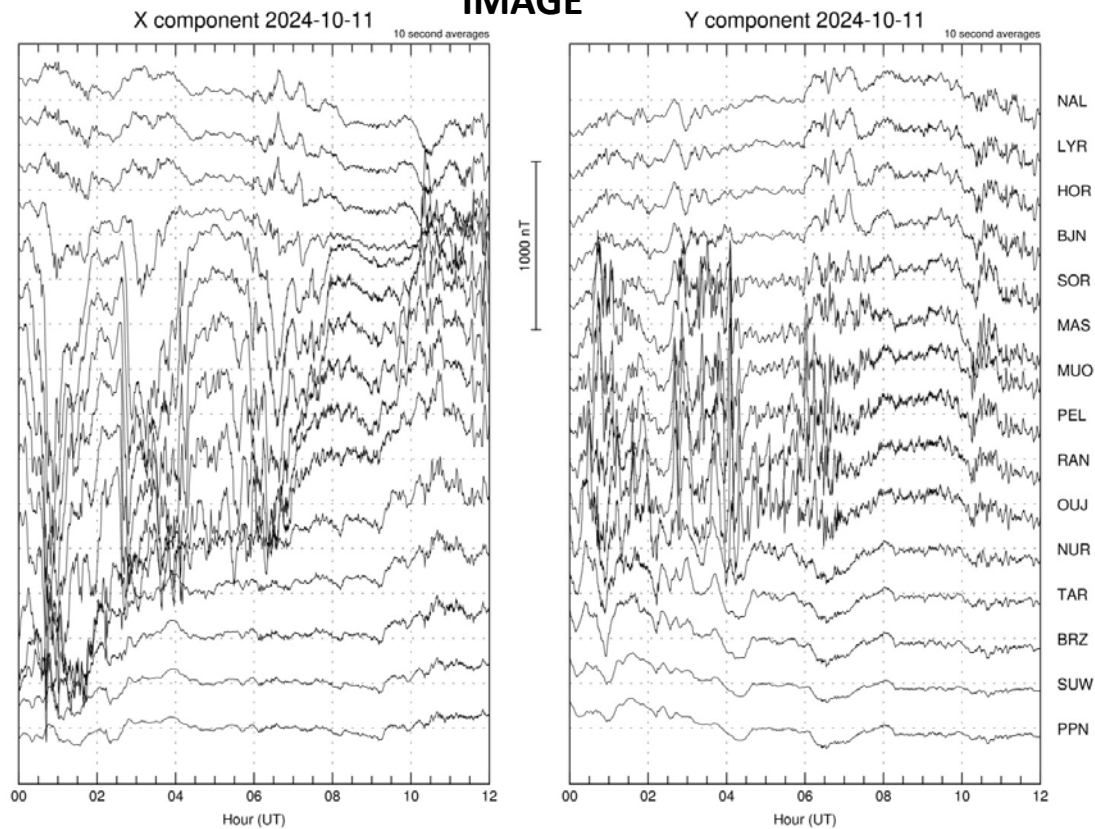
MIRACLE



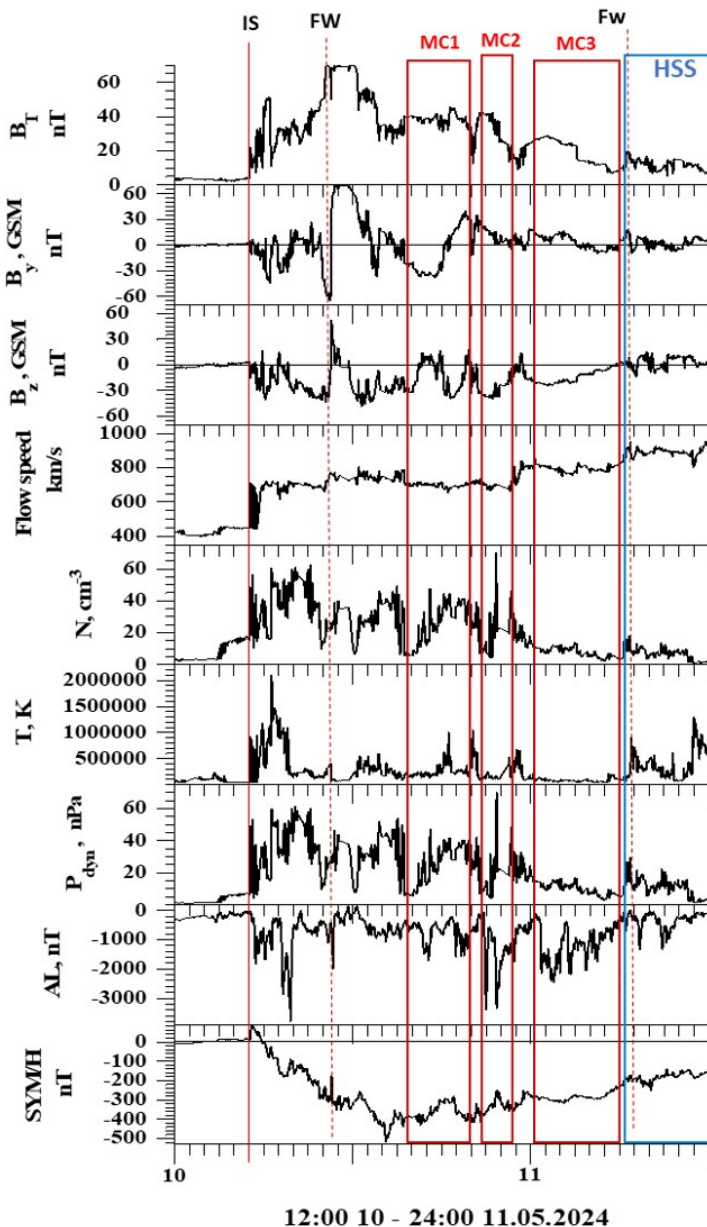
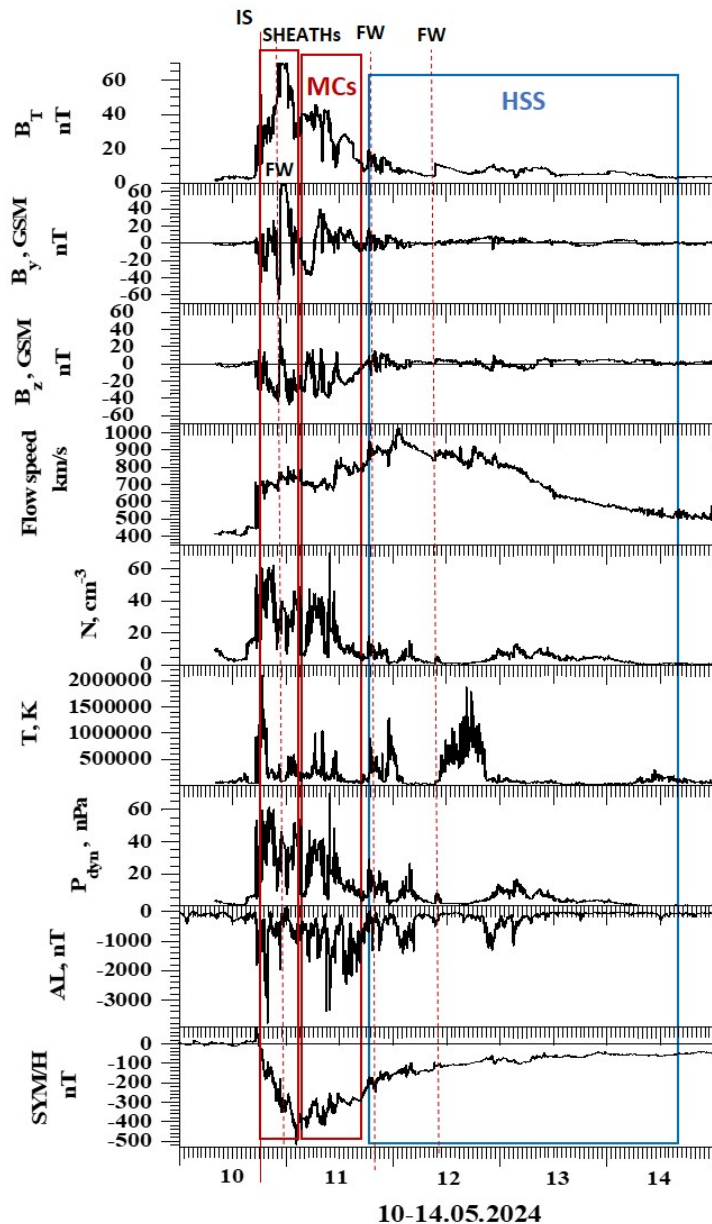
IMAGE+GIC



IMAGE



Solar wind/interplanetary and geomagnetic conditions during the 2024 May, 10-12 superstorm (SYM/H~-518 nT)



Interplanetary shock wave arrived at L1 at 16:45 UT on May 10 (Gonzales-Esparza et al. 2024; Hajra et al. 2024).

three-stage strong magnetic storm :
SYM/H ~-183nT;
-354 nT;
-518 nT

MCs

- 1) ~03:28 to 07:54 UT,
 - 2) ~08:25 to 10:36 UT,
 - 3) ~11:19 to 17:17 UT
- on May 11

GICs and geomagnetic indexes registered in 10- 13 May 2024

Night sector, 10 May :

~18:30 UT, ~62A (80A) (KND)
~20:20 UT, ~26A (KND)
~22:10 UT, ~25A (KND)

Morning and day
sector, 11 May:

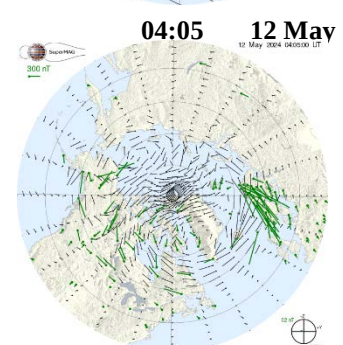
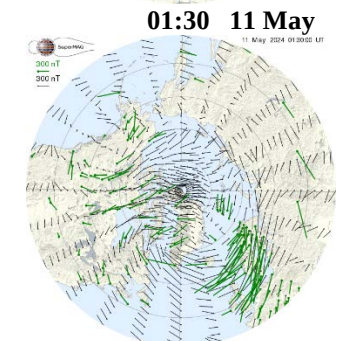
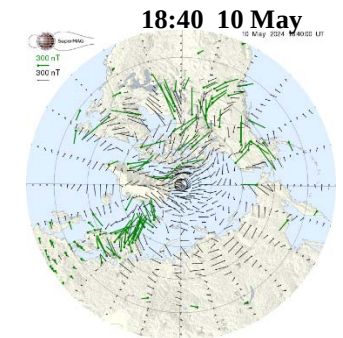
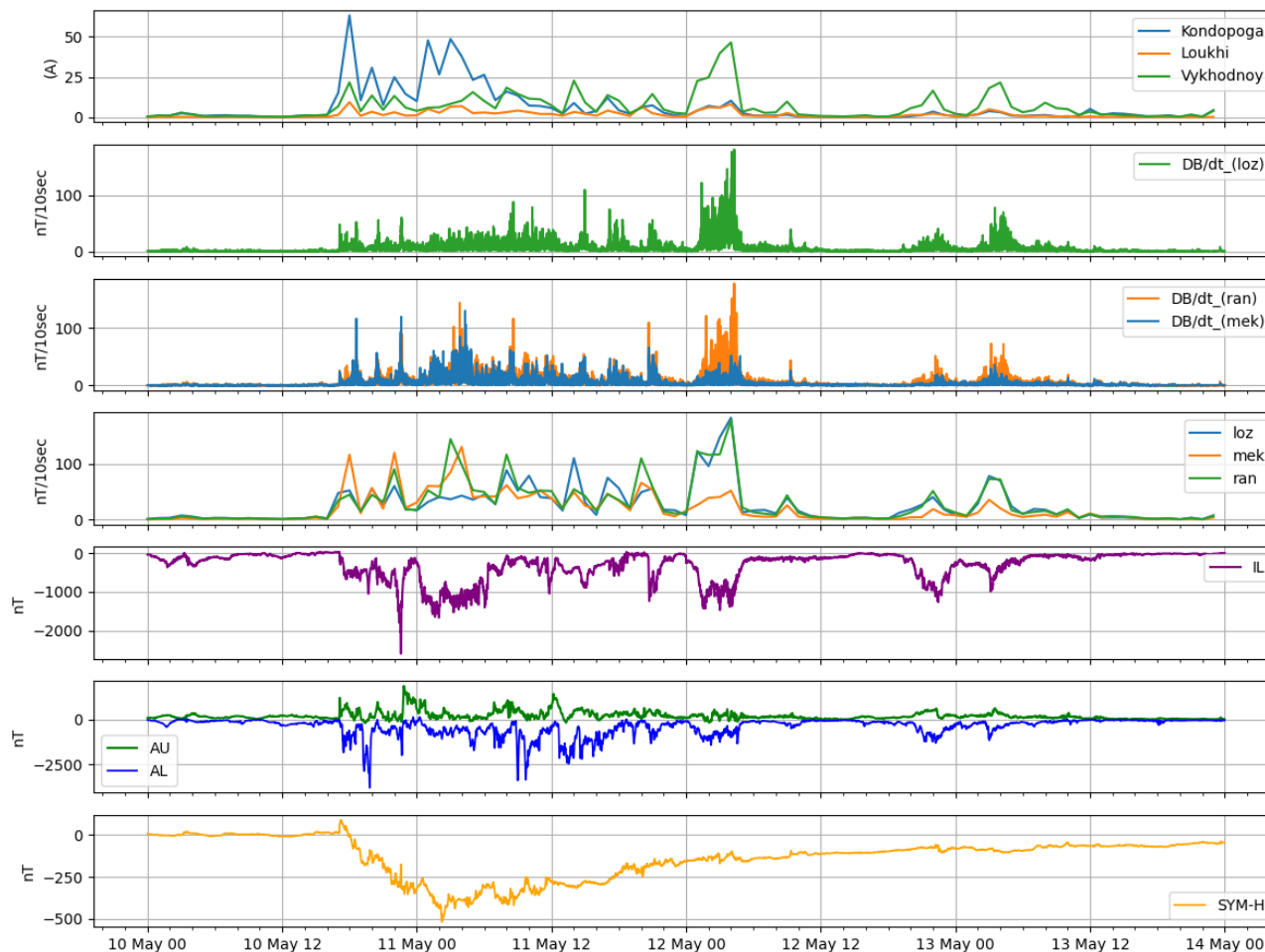
~01:40 UT, ~50A (KND)
~03:20 UT, ~50A (KND)
~08:40 UT, ~20A (VKH)

Evening sector,
11 May:

~15:00 UT, ~22A (VKH)
~17:00 UT, ~12A (VKH)
~20:40 UT, ~12A (VKH)

Morning sector,
12 May:

~01:30 UT, ~20A (VKH)
~01:50 UT, ~25A (VKH)
~04:00 UT, ~50A (VKH)



Evening and pre-midnight sectors on 10 May

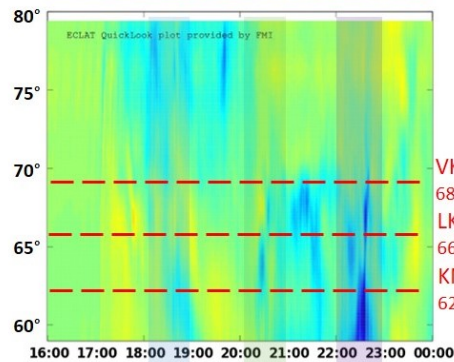
- GIC sources in the evening sector:
the superposition of the substorm development and a local magnetic disturbance caused by a large solar wind pressure jump.

MIRACLE

10 May 2024

b) E (+) and W (-) electrojet at 22.06° lon

A/km

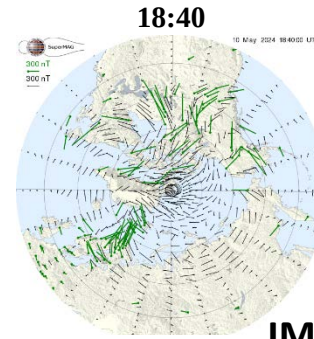
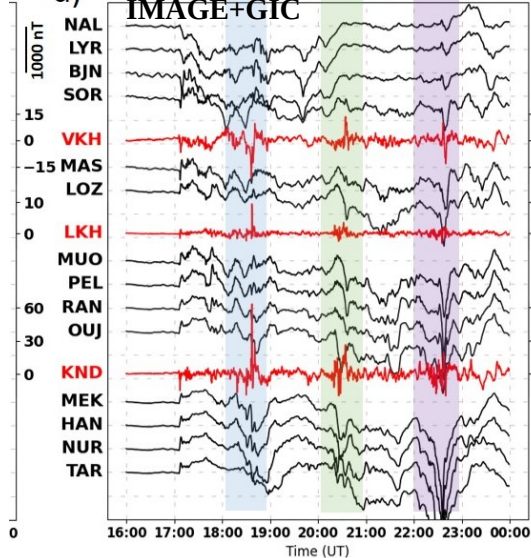


VKH
68.8° N
LKH
66.1° N
KND
62.2° N

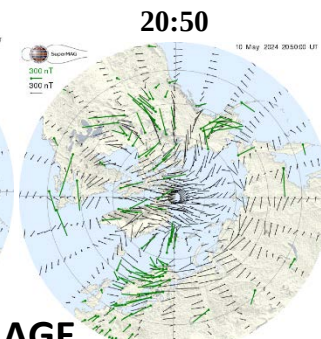
1000 nT

d) IMAGE X component

IMAGE+GIC



18:40



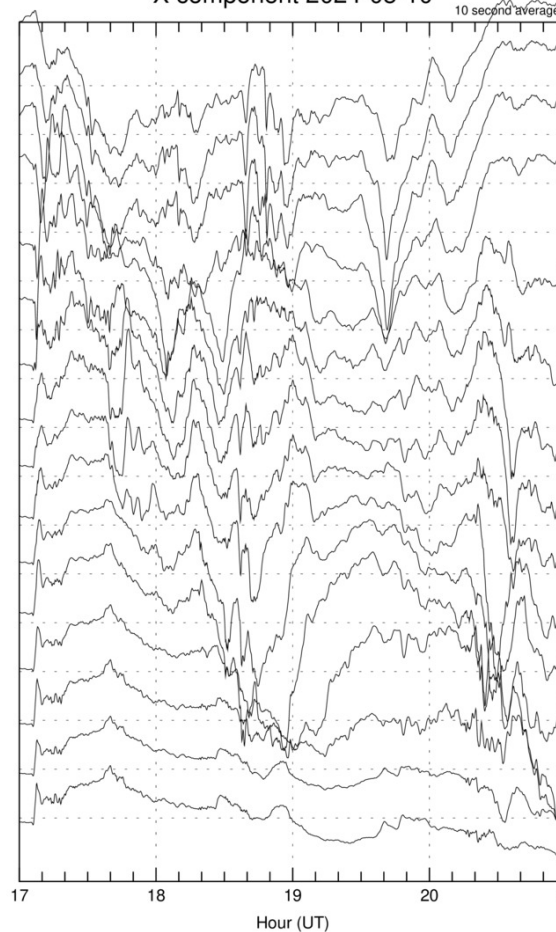
20:50

SuperMAG

IMAGE

X component 2024-05-10

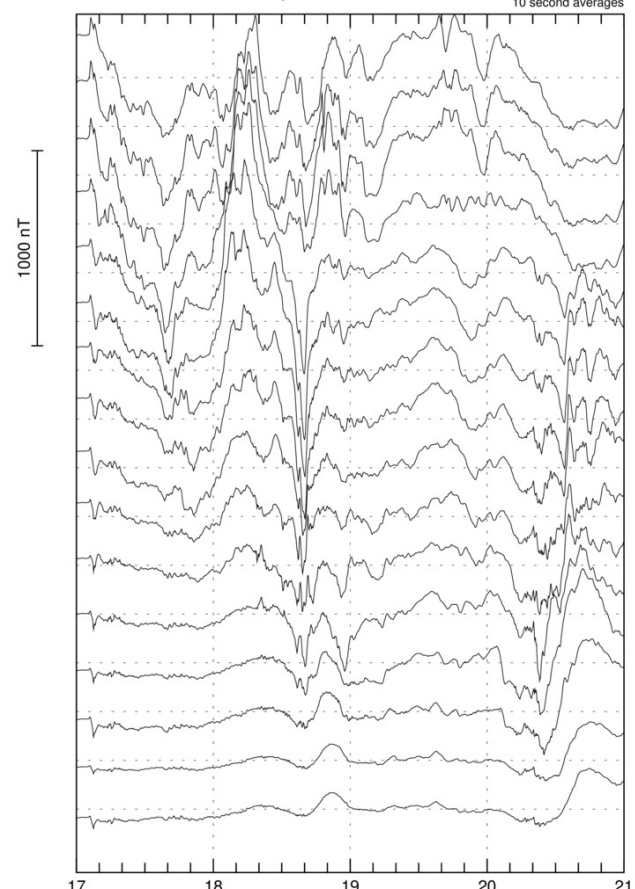
10 second averages



Hour (UT)

Y component 2024-05-10

10 second averages

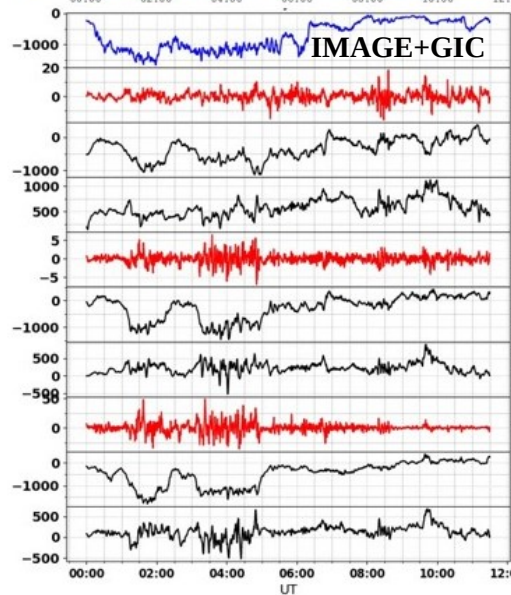
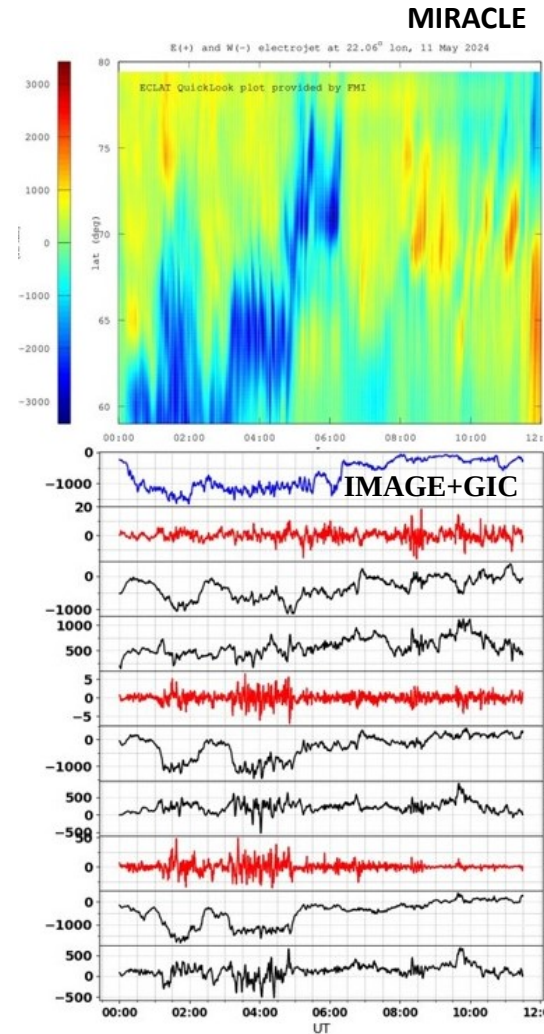


Hour (UT)

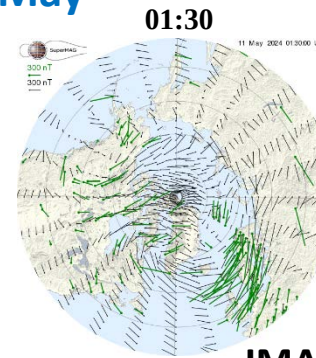
NAL
LYR
HOR
BJN
SOR
MAS
MUO
PEL
RAN
OUJ
HAN
NUR
TAR
BRZ
SUW
PPN

Morning interval on 11 May

- GIC sources in the morning sector:
Pc5/Pi3 geomagnetic pulsations and
substorms

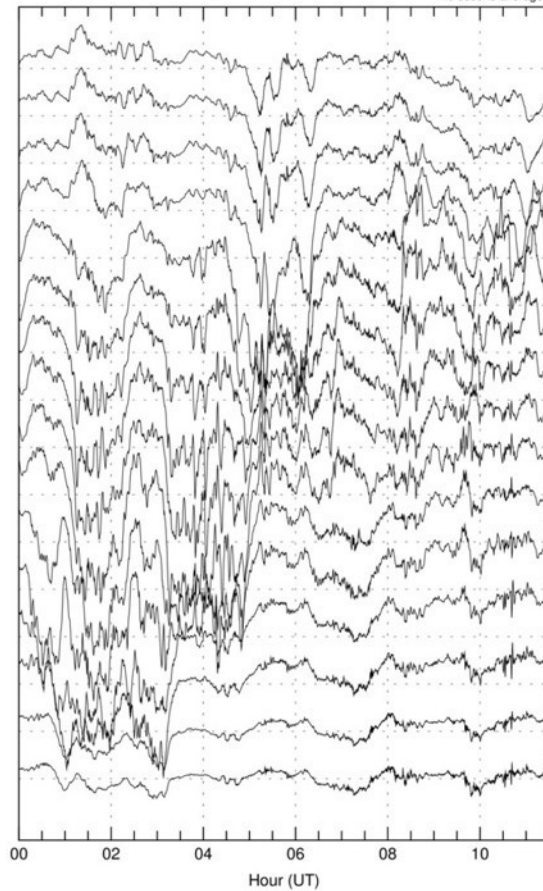


SuperMAG

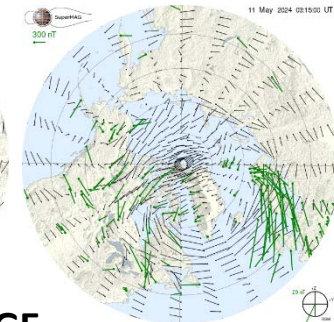


IMAGE

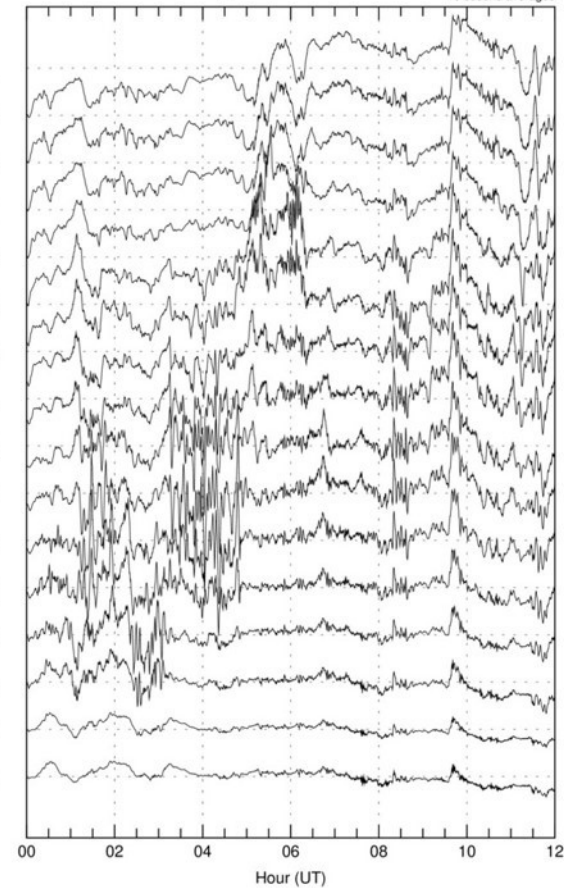
X component 2024-05-11



03:15

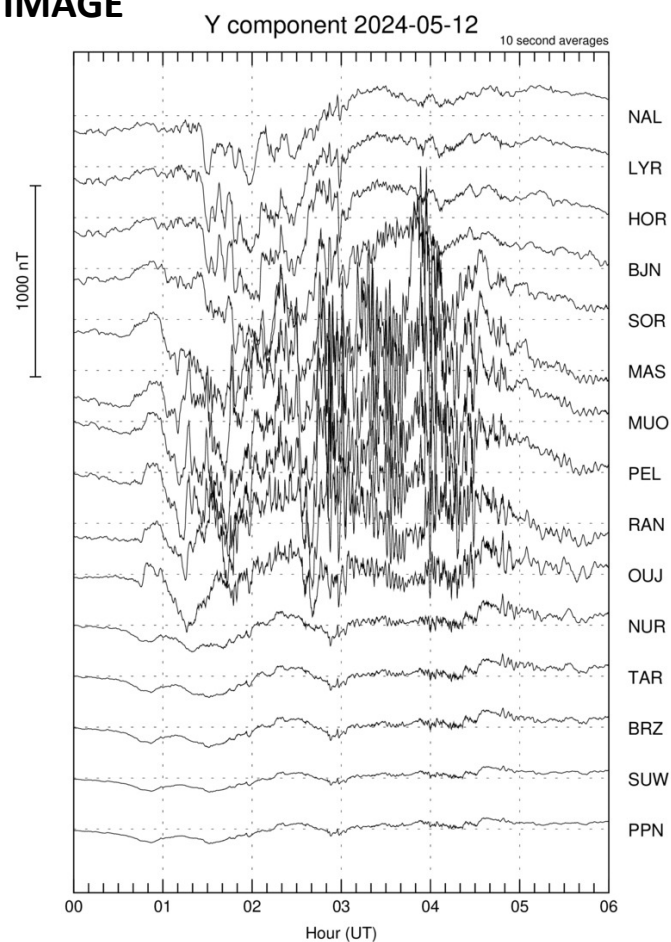
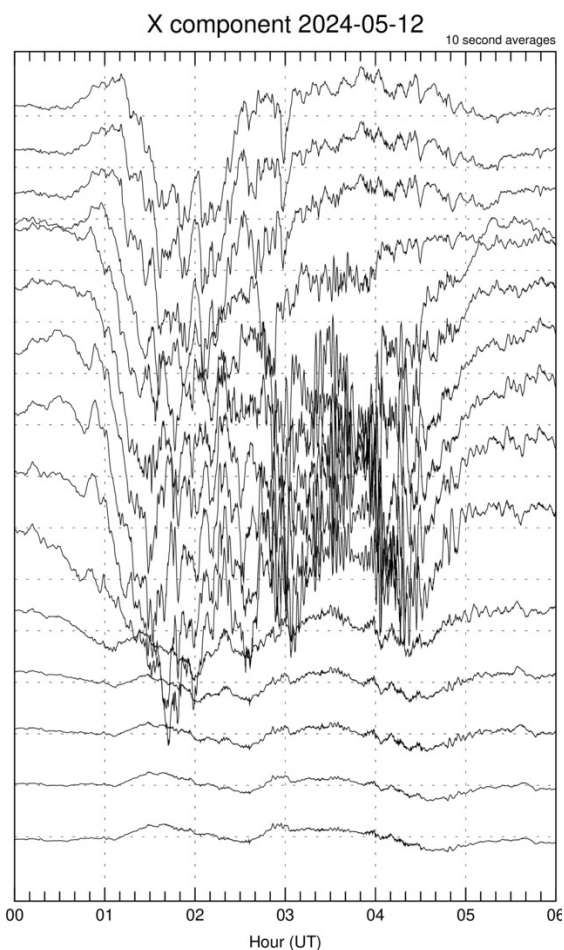
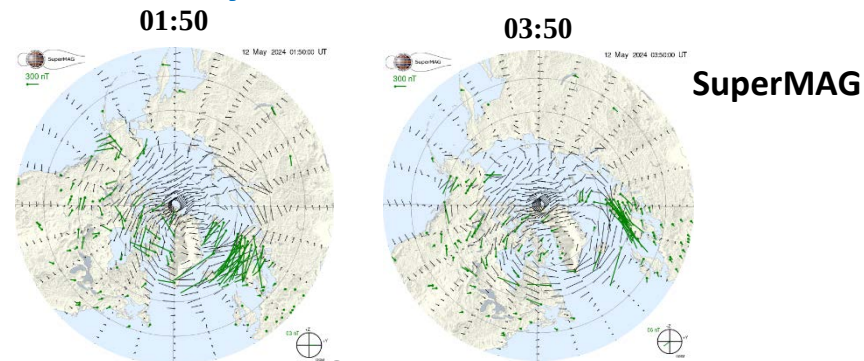
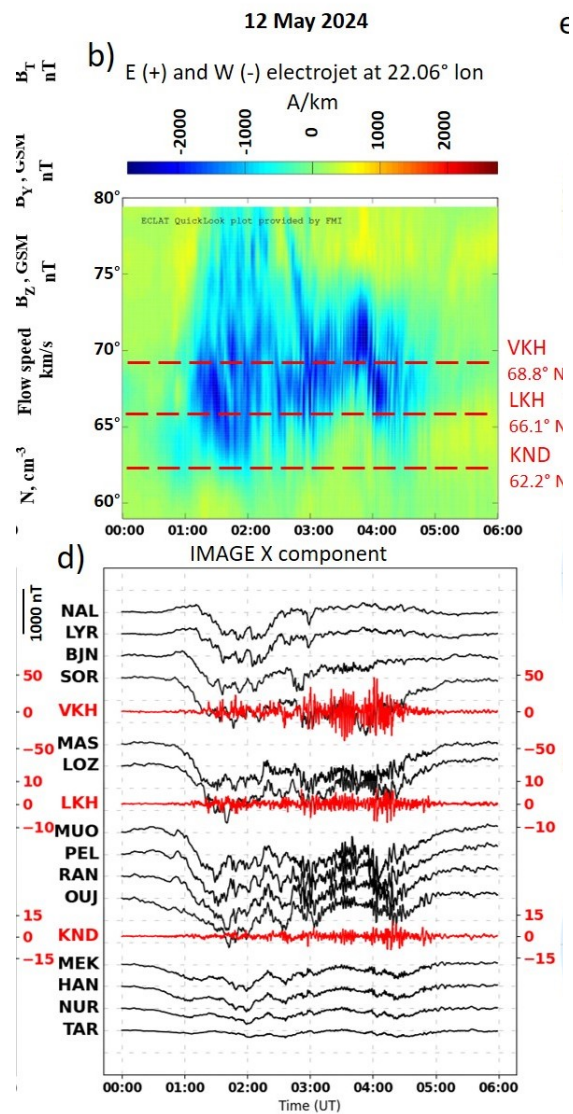


Y component 2024-05-11



Morning interval on 12 May

- GIC sources in the morning sector:
- the superposition of Pc5/Pi3 geomagnetic pulsations and the substorm development



Conclusions

GIC sources are distributed over MLT sectors:

an intense GICs (**~10–30A**) in the evening and night sectors were caused by **substorms**, while intense GICs in the morning sectors were caused by **Pc5/Pi3 geomagnetic pulsations**.

One strongest GIC peak was recorded during storm on **10-11 October 2024**:

- **GIC peak of ~30 A** at **VKH** station was ~ 15:20 UT at the beginning of the magnetic storm and was probably associated with the **IS** or **SC**.

Three strongest GIC peaks were identified during superstorm on **10-12 May 2024**:

Three strongest GIC peaks of **~50-62 A** - one in the night and two in the morning sectors- were associated with the superposition of two different geophysical sources:

- First strongest GIC (**~62A**) registered at ~18:40 UT on 10 May at **KND** was caused the superposition of the substorm development and a local magnetic disturbance caused by a large solar wind pressure jump.
- Second strongest GIC peaks (**~50 A**) at ~03:30 UT on May 11 at **KND** were associated with the superposition of Pc5/Pi3 geomagnetic pulsations and negative bays of subsequent substorms.
- Third strongest GIC peak (**~50 A**) at ~04:00 UT on May 12 at **VKH** was associated with strong Pc5/Pi3 geomagnetic pulsations recorded at the recovery phase of substorm recorded during HSS.